ACCURATE SOLUTIONS FOR RADIATIVE HEAT TRANSFER IN TWO-DIMENSIONAL AXISYMMETRIC ENCLOSURES WITH GAS RADIATION AND REFLECTIVE SURFACES

P. PEREZ[*]**, M. EL HAFI**[*]**, P. J. COELHO**[**]**, R. FOURNIER**[***]

[*] École des Mines d'Albi-Carmaux

Campus Jarlard, Route de Teillet, 81013 Albi CT Cedex 09

France

[**] Instituto Superior Técnico

Mechanical Engineering Department

Av. Rovisco Pais, 1049-001 Lisboa

Portugal

[***] Laboratoire d'Énergétique

U. Paul Sabatier, Route de Narbonne, 31000 Toulouse

France

Corresponding author: P. Perez

Phone:(+33)5 63 49 32 49

Fax:(+33)5 63 49 32 43

E-mail: patperez@enstimac.fr

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Abstract

Accurate solutions for benchmarking purposes in two-dimensional axisymmetric enclosures with reflective surfaces have been obtained using the Monte Carlo method (MCM) based on the net exchange formulation (NEF). Previous applications of the MCM-NEF have been restricted to multidimensional problems with black boundaries or one-dimensional problems with grey boundaries. Here, the extension to multidimensional enclosures with grey boundaries is presented. The medium is a mixture of H_2O, CO_2, N_2 and soot at atmospheric pressure, and its radiative properties are computed using the correlated k-distribution method. Predictions obtained using the discrete ordinates method are included, showing a good agreement with the benchmark MCM/NEF solutions.

Nomenclature

A area of cell faces (DOM)

f_v soot volumetric fraction

$\mathcal{F}$ directional probability density function for reflexion

g cumulative distribution function of the absorption coefficient

I_η spectral radiation intensity

L length

M number of directions

N_b number of narrow bands

N_Q number of quadrature points (CK method)

N_s number of surfaces

N_v number of volumes

p_n probability of band number n

$\boldsymbol{q}$ heat flux vector

r radial coordinate

$\boldsymbol{r}$ position vector

R radius

S surface

$\boldsymbol{u}$ direction vector

V volume or cell volume

w_m DOM quadrature weight for direction m

w_k DOM quadrature weight for CK method

W Monte Carlo integration weight

x axial coordinate

Greek symbols

α	absorptivity
α_m	curvature coefficient of angular redistribution (DOM)
β	direction cosine
Γ	visibility factor
ϵ	accuracy
ε	emissivity
δ	Dirac function
η	wave number
θ	angle between direction $\boldsymbol{u}$ and normal to a surface
κ	absorption coefficient
μ	direction cosine
ξ	direction cosine
σ	length
τ_{ij}	spectral transmissivity from point P_i to point P_j
$\varphi(S_i, S_j)$	net radiative exchange between surfaces S_i and S_j
$\varphi(S_i, V_j)$	net radiative exchange between surface S_i and volume V_j
$\varphi(V_i, S_j)$	net radiative exchange between volume V_i and surface S_j
$\varphi(V_i, V_j)$	net radiative exchange between volumes V_i and V_j
ω	solid angle

Subscripts

w	wall

Superscripts

m	direction
0	blackbody
—	mean value

1 Introduction

In these last decades, numerous authors have pointed out the importance of radiative heat transfer in combustion systems such as diesel engines, boilers, furnaces, rocket engines and many other practical applications which involve high temperatures and large scales [1, 2]. Only few analytical or quasi-exact solutions are available in the literature for simple geometries such as two-dimensional rectangular/axisymmetrical and three-dimensional rectangular enclosures with gray media. Accurate results of multi-dimensional non-gray radiation analysis of real gases is typically lacking. That is mainly due to the unacceptable computing time required by the detailed models of gaseous radiation calculations in multi-dimensions. Considering this state of the art, a first benchmark numerical solution has been proposed recently by Coelho et al. [3], for two-dimensional black enclosures with non-gray sooting media. The aim of the present paper is to extend this work to enclosures bounded by gray diffuse and specular walls. When using the MCM with a net exchange formulation, it is not possible any more to rely on the physical pictures of photon statistical transport to derive the algorithm. It is required that the algorithm is rigorously justified by a complete formulation work. This was only done with reflexion in [4] for 1D geometries in a way that could not be simply extended to complex geometries. It is therefore the first time that the extension of Net-Exchange MCM to reflexion in any geometry is presented for publication. A new formulation of the MCM-NEF (Monte Carlo-Net Exchange Formulation) involving diffuse and specular walls is detailed in the section devoted to the theoretical methods. The results presented in the last section are useful in evaluating the accuracy of other approximate numerical methods such as the DOM (Discrete Ordinates method) which remains adequate for most practical applications [5, 6, 7].

2 Theoretical methods

2.1 The Net Exchange Formulation and the Monte Carlo Method.

The many variants of the Monte Carlo method (MCM) are generally recognized as accurate solution method (see for instance the review of J.R. Howell in [8]), and are commonly used for benchmark purposes. Most of the techniques are presented as strict numerical implementations of photon transport stochastic models and optimization efforts essentially concern the adaptation of the random sampling procedures [9, 10]. However, these techniques based on standard photon transport formulations frequently appeared to be not efficient and to require very large numbers of statistical realizations for accurate solutions, when optical thicknesses are high, when configurations include hot spots, or at the opposite when temperature differences are small. Recently, numerous research efforts have been undertaken to overcome these difficulties on the basis of optimized formulations [8, 11, 12, 13, 14, 4, 15, 16, 17]. Among such optimization techniques, the MCM-NEF was shown to have many advantages in terms of statistical convergence, but a detailed formulation procedure is strictly required for each extension to new application fields [4, 16, 18]. The Net Exchange Formulation (NEF) was originally proposed in [19] for monochromatic radiative exchanges in the atmosphere, and later extended in [20] to narrow bands. It has some similarities with the zone method, but it does not require that the surface and volume zones are isothermal. In the field of thermal radiation, the Monte Carlo - Net Exchange Formulation method was firstly applied to one-dimensional enclosures in [14]. In this method, the net radiative heat exchanges between every volume-volume, surface-volume and surface-surface pair are calculated using the Monte Carlo method for the computation of the integrals that appear in the definition of the net radiative exchanges. One of the advantages of the MCM-NEF over the traditional MCM is that it intrinsically satisfies the reciprocity principle and the energy conservation, which allows to bypass convergence difficulties in the quasi-isothermal limit and in the optically thick limit [15]. The method was extended to two-dimensional Cartesian enclosures with black walls in [16] and to one-dimensional enclosures with reflective surfaces in [4].A very recent work using this methodology to address optically thick scattering was carried out by Eymet and al. in [18, 21]. Theoretical

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derivations are presented in sections 2.1.1, 2.1.2 and 2.1.3 that allow MCM-NEF applications to multi-dimensional enclosures with reflective surfaces.

2.1.1 The mathematical formulation.

Let P_i be a point within volume V_i, and Q_1 be a point on surface S_i defined respectively by the position vectors $\boldsymbol{r_{P_i}}$ and $\boldsymbol{r_{Q_1}}$ (Figure 1). The net radiative exchange between two volumes V_i and V_j, $\varphi_{(V_i,V_j)}$, or a volume V_i and a surface S_j, $\varphi_{(V_i,S_j)}$, or a surface S_i and a volume V_j, $\varphi_{(S_i,V_j)}$, or a surface S_i and a surface S_j, $\varphi_{(S_i,S_j)}$, are expressed as follows for black walls and non scattering media:

$$\varphi_{(V_i,V_j)} = \int_0^\infty d\eta \int_{V_i} dV_i \int_{V_j} dV_j \frac{\kappa_i \tau_{ij} \kappa_j}{\sigma_{ij}^2} \Delta I_{ij} \tag{1}$$

$$\varphi_{(V_i,S_j)} = \int_0^\infty d\eta \int_{V_i} dV_i \int_{S_j} dS_j \frac{\kappa_i \tau_{ij}}{\sigma_{ij}^2} \Delta I_{ij} \tag{2}$$

$$\varphi_{(S_i,V_j)} = \int_0^\infty d\eta \int_{S_i} dS_i \int_{V_j} dV_j \frac{\cos\theta_i \tau_{ij} \kappa_j}{\sigma_{ij}^2} \Delta I_{ij} \tag{3}$$

$$\varphi_{(S_i,S_j)} = \int_0^\infty d\eta \int_{S_i} dS_i \int_{S_j} dS_j \frac{\cos\theta_i \tau_{ij}}{\sigma_{ij}^2} \Delta I_{ij} \tag{4}$$

A shortened notation is adopted for the optical properties $\kappa_i, \kappa_j, \tau_{ij}$ where $[\kappa_i \tau_{ij} \kappa_j]$ means $[\kappa_\eta(\boldsymbol{r_{P_i}}) \tau_\eta(\boldsymbol{r_{P_i}} \rightarrow \boldsymbol{r_{P_j}}) \kappa_\eta(\boldsymbol{r_{P_j}})]$. The Planck function ΔI_{ij} here is the difference of the black body function taken at each point (for instance : $\Delta I_{ij} = I_\eta^0(\boldsymbol{r_{P_i}}) - I_\eta^0(\boldsymbol{r_{P_j}})$). σ_{ij} is the distance between point P_i and point P_j.

Considering two volumes V_i and V_j, the net radiative exchange between them is expressed as a multiple integral on the wave number η and on the considered volumes. In the Monte Carlo method, the integrations are performed by generating a large number of rays, and each ray is defined by a frequency, a first exchange point P_i, and a direction $\boldsymbol{u_0}$ supporting a solid angle ω_0. This is accomplished in terms of formulation by simple variable substitutions transforming for instance a volume integration over V_j into a double integral over solid angle and over the intersecting segment (see Fig. 1).

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$$\varphi_{(V_i,V_j)} = \int_0^\infty d\eta \int_{V_i} dV_i \int_{4\pi} d\omega_0 \left[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}},\boldsymbol{u_0})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} \right] \tag{5}$$

$$\varphi_{(V_i,S_j)} = \int_0^\infty d\eta \int_{V_i} dV_i \int_{4\pi} d\omega_0 \left[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \kappa_i \tau_{ij} \Delta I_{ij} \right] \tag{6}$$

$$\varphi_{(S_i,V_j)} = \int_0^\infty d\eta \int_{S_i} dS_i \int_{2\pi} d\omega_0 \left[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}},\boldsymbol{u})} d\sigma \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} \right] \tag{7}$$

$$\varphi_{(S_i,S_j)} = \int_0^\infty d\eta \int_{S_i} dS_i \int_{2\pi} d\omega_0 \left[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \cos\theta_i \tau_{ij} \Delta I_{ij} \right] \tag{8}$$

For all the exchanges : $\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) = 1$ if the optical path intersects the exchange area j (surface S_j or volume V_j) and $\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) = 0$ if not. At this step, still considering an exchange between V_i and V_j, the volume V_j may be intercepted by the generated ray, defining a segment $\sigma_0(\boldsymbol{r_{P_i}}, \boldsymbol{u_0})$ inside V_j. The second exchange point is then chosen on this segment at curvilinear abscissa σ and the exchange is able to be estimated for this ray.

The net radiative heat flux on a surface S_i is then obtained by summing all radiative exchanges between surface S_i and all other surfaces and volumes of the system:

$$q_{w,\text{net},i} = \sum_{j=1}^{N_v} \varphi_{(S_i,V_j)} + \sum_{j=1}^{N_s} \varphi_{(S_i,S_j)} \tag{9}$$

The radiative heat source for volume V_i is similarly:

$$\int_{V_i} \nabla.\boldsymbol{q} dV_i = \sum_{j=1}^{N_v} \varphi_{(V_i,V_j)} + \sum_{j=1}^{N_s} \varphi_{(V_i,S_j)} \tag{10}$$

2.1.2 Generalization of the NEF to multiple reflexions.

Multiple reflexions are now considered in the NEF. New terms appear compared with equations 5,6,7, and 8 in order to take into account the contribution of multiple reflexion rays in the net exchanges (Figure 2).

Exchange between two volumes

The net radiative exchange between two volumes is given by:

$$
\begin{aligned}
\varphi_{(V_i,V_j)} &= \int_0^\infty d\eta \int_{V_i} dV_i \int_{4\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r}_{\boldsymbol{P}_i}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r}_{\boldsymbol{P}_i}, \boldsymbol{u_0})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} \\
&+ \int_{2\pi} d\omega_1 \Bigg[\Gamma_j(\boldsymbol{r}_{\boldsymbol{Q_1}}, \boldsymbol{u_1}) \int_{\sigma_1(\boldsymbol{r}_{\boldsymbol{Q_1}}, \boldsymbol{u_1})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&+ \int_{2\pi} d\omega_2 \Big[\Gamma_j(\boldsymbol{r}_{\boldsymbol{Q_2}}, \boldsymbol{u_2}) \int_{\sigma_2(\boldsymbol{r}_{\boldsymbol{Q_2}}, \boldsymbol{u_2})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&\times \quad (1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \Big] \Bigg] \Bigg]
\end{aligned}
\tag{11}
$$

$\varepsilon_1, \varepsilon_2, \ldots$, stands for the directional emissivities at the reflexion points in directions $-\boldsymbol{u_0}, -\boldsymbol{u_1}$, $\ldots$ (or the directional absorptivities in directions $+\boldsymbol{u_0}, +\boldsymbol{u_1}, \ldots$): $\varepsilon_1 = \varepsilon(\boldsymbol{r}_{\boldsymbol{Q_1}}, -\boldsymbol{u_0}) = \alpha(\boldsymbol{r}_{\boldsymbol{Q_1}}, \boldsymbol{u_0})$, $\varepsilon_2 = \varepsilon(\boldsymbol{r}_{\boldsymbol{Q_2}}, -\boldsymbol{u_1}) = \alpha(\boldsymbol{r}_{\boldsymbol{Q_2}}, \boldsymbol{u_1})$.

$\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})$, $\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})$ are the reflexion phase functions at the successive reflexion points ($\int_{2\pi} \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) d\omega_1 = 1$). For a diffuse reflexion, $\varepsilon_1(\boldsymbol{u_0}) = \varepsilon_1$ (independent of direction) and $\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) = (\boldsymbol{u_1}.\boldsymbol{n}(\boldsymbol{r}_{\boldsymbol{Q_1}}))/\pi$, and for a specular reflexion, $\varepsilon_1(\boldsymbol{u_0}) = \varepsilon_1$ and $\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) = \delta(\boldsymbol{u_1} - sym(-\boldsymbol{u_0}, \boldsymbol{r}_{\boldsymbol{Q_1}}))$ where δ is the Dirac function and $sym(-\boldsymbol{u_0}, \boldsymbol{r}_{\boldsymbol{Q_1}})$ is the direction symmetric to $-\boldsymbol{u_0}$ relatively to the outer normal vector of the surface.

In equation 11, the first line is similar to the expression of $\varphi_{(V_i,V_j)}$ for black walls (equation 5), and the contribution of the reflexions appear on the following lines of the equation. Physically, after the first reflexion at point Q_1, the ray propagates in a new direction $\boldsymbol{u_1}$, and may once again intercept the volume V_j defining inside a new segment $\sigma_1(\boldsymbol{r}_{\boldsymbol{Q_1}}, \boldsymbol{u_1})$. The second line of equation 11 expresses the contribution of this new path to the net exchange between V_i and V_j. Reasoning in term of NEF on the basis of this physical picture only requires to consider each defined optical path as traveled both way from V_i to V_j and from V_j to V_i [4].

Altogether, the first line represents the contribution to $\varphi_{(V_i,V_j)}$ of all the optical paths between V_i and V_j without reflexion. The second line represents the contribution of all the optical paths after one reflexion, the third line stands for the contribution after two reflexions, and so on.

The subscript j is employed in κ_j, τ_{ij} and ΔI_{ij} for all the terms of equation 11, in each integral, but the reader should be aware that the meanings are different after each reflexion because the

corresponding variables are taken at different points in different optical paths.

Exchange between a volume and a surface

Physical considerations are the same as for the exchange between two volumes, except that integration over $\sigma_0,\sigma_1,\sigma_2$ (through the volume) is not required : exchanges only occur with surface points. The first line of equation (12) is similar to the equation without reflexion (equation 6) and supplementary contributions are expressed in the following lines, representing respectively the energy exchanged between V_i and S_j after one, two, ... reflexions.

$$
\begin{aligned}
\varphi_{(V_i,S_j)} &= \int_0^\infty d\eta \int_{V_i} dV_i \int_{4\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0})\kappa_i\tau_{ij}\varepsilon_1\Delta I_{ij} \\
&+ \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1})\kappa_i\tau_{ij}\varepsilon_2\Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&+ \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2})\kappa_i\tau_{ij}\varepsilon_3\Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&\times \ (1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{12}
$$

The exchanges between a volume and a surface (equation 13), and between two surfaces (equation 14) given hereafter are obtained in the same way.

Exchange between a surface and a volume

$$
\begin{aligned}
\varphi_{(S_i,V_j)} &= \int_0^\infty d\eta \int_{S_i} dS_i \int_{2\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}},\boldsymbol{u_0})} d\sigma\varepsilon_i \cos\theta_i\tau_{ij}\kappa_j\Delta I_{ij} \\
&+ \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1}) \int_{\sigma_1(\boldsymbol{r_{Q_1}},\boldsymbol{u_1})} d\sigma\varepsilon_i \cos\theta_i\tau_{ij}\kappa_j\Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&+ \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2}) \int_{\sigma_2(\boldsymbol{r_{Q_2}},\boldsymbol{u_2})} d\sigma\varepsilon_i \cos\theta_i\tau_{ij}\kappa_j\Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
&\times \ (1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{13}
$$

Exchange between two surfaces

$$
\begin{aligned}
\varphi_{(S_i,S_j)} \;=\; & \int_0^\infty d\eta \int_{S_i} dS_i \int_{2\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_1 \Delta I_{ij} \\
& + \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_2 \Delta I_{ij} (1-\varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& + \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_3 \Delta I_{ij} (1-\varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& \times \; (1-\varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{14}
$$

The wave number integration is carried out over narrow-bands $\Delta\eta$ according to the Statistical Narrow Band model proposed by Malkmus [22], using a k-distribution reformulation [23] and a correlated-k assumption for the representation of inhomogeneities [24]. In the correlated-k distribution approach, frequencies are rearranged and the absorption spectrum κ_η is replaced with an equivalent spectrum κ_g in which κ is a monotonous function of a pseudo frequency g. Practically g is the cumulative of the k-distribution and each frequential integration over a narrow band $\Delta\eta$ is transformed as $\int_{\Delta\eta} d\eta \ldots = \Delta\eta \int_0^1 dg \ldots$. The shortened notation $\kappa_i \tau_{ij} \kappa_j$, for instance, now stands for $\kappa_g(\boldsymbol{r_{P_i}}) \tau_g(\boldsymbol{r_{P_i}} \to \boldsymbol{r_{P_j}}) \kappa_g(\boldsymbol{r_{P_j}})$. Then, from equations 11, 12, 13, 14, it yields:

$$
\begin{aligned}
\varphi_{(V_i,V_j)} \;=\; & \sum_{n=1}^{N_b} \Delta\eta_n \int_0^1 dg \int_{V_i} dV_i \int_{4\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}}, \boldsymbol{u_0})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} \\
& + \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1}) \int_{\sigma_1(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} (1-\varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& + \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2}) \int_{\sigma_2(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2})} d\sigma \kappa_i \tau_{ij} \kappa_j \Delta I_{ij} (1-\varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& \times \; (1-\varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{15}
$$

Where N_b is the total number of narrow bands. The parameters of the Malkmus model were taken from the database given in [25] for 367 narrow bands of $25cm^{-1}$ width considering a spectral range from $150cm^{-1}$ to $9300cm^{-1}$. The number of narrow bands has been extended to 800 in order to take into account the whole spectral range from 0 to $20000cm^{-1}$ for soot radiation at high frequencies for high temperatures.

$$
\begin{aligned}
\varphi_{(V_i, S_j)} \;=\; & \sum_{n=1}^{N_b} \Delta\eta_n \int_0^1 dg \int_{V_i} dV_i \int_{4\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \kappa_i \tau_{ij} \varepsilon_1 \Delta I_{ij} \\
& + \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{Q_1}, \boldsymbol{u_1}) \kappa_i \tau_{ij} \varepsilon_2 \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& + \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2}) \kappa_i \tau_{ij} \varepsilon_3 \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& \times (1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{16}
$$

$$
\begin{aligned}
\varphi_{(S_i, V_j)} \;=\; & \sum_{n=1}^{N_b} \Delta\eta_n \int_0^1 dg \int_{S_i} dS_i \int_{2\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}}, \boldsymbol{u})} d\sigma \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} \\
& + \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1}) \int_{\sigma_1(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1})} d\sigma \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& + \int_{2\pi} d\omega_2 \big[\Gamma_j(r_{Q_2}, \boldsymbol{u_2}) \int_{\sigma_2(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2})} d\sigma \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& \times (1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{17}
$$

$$
\begin{aligned}
\varphi_{(S_i, S_j)} \;=\; & \sum_{n=1}^{N_b} \Delta\eta_n \int_0^1 dg \int_{S_i} dS_i \int_{2\pi} d\omega_0 \Bigg[\Gamma_j(\boldsymbol{r_{P_i}}, \boldsymbol{u_0}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_1 \Delta I_{ij} \\
& + \int_{2\pi} d\omega_1 \Big[\Gamma_j(\boldsymbol{r_{Q_1}}, \boldsymbol{u_1}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_2 \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& + \int_{2\pi} d\omega_2 \big[\Gamma_j(\boldsymbol{r_{Q_2}}, \boldsymbol{u_2}) \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_3 \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0}) \\
& \times (1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \big] \Big] \Bigg]
\end{aligned}
\tag{18}
$$

2.1.3 Monte Carlo integrations

A Monte Carlo algorithm based on this formulation has been developed with the use of adapted probability density functions (pdf) in order to optimize the number of sampling events and to ensure a fast convergence, especially for optically thick media [15]. For instance, if pdf functions are introduced in the net exchange between two volumes (equation 15), one obtains:

$$\varphi_{(V_i,V_j)} = \sum_{n=1}^{N_b} \Delta\eta_n p_n \frac{1}{p_n} \int_0^1 pdf(g)\frac{1}{pdf(g)}dg \int_{V_i} pdf(V_i)\frac{1}{pdf(V_i)}dV_i$$

$$\int_{4\pi} pdf(\omega_0)\frac{1}{pdf(\omega_0)}d\omega_0 \left[\Gamma_j(\boldsymbol{r_{P_i}},\boldsymbol{u_0}) \int_{\sigma_0(\boldsymbol{r_{P_i}},\boldsymbol{u_0})} pdf(\sigma_0)\frac{1}{pdf(\sigma_0)}d\sigma_0\right.$$

$$\times \quad \kappa_i \tau_{ij} \kappa_j \Delta I_{ij}$$

$$+ \quad \int_{2\pi} pdf(\omega_1)\frac{1}{pdf(\omega_1)}d\omega_1 \left[\Gamma_j(\boldsymbol{r_{Q_1}},\boldsymbol{u_1}) \int_{\sigma_1(\boldsymbol{r_{Q_1}},\boldsymbol{u_1})} pdf(\sigma_1)\frac{1}{pdf(\sigma_1)}d\sigma_1\right.$$

$$\times \quad \kappa_i \tau_{ij} \kappa_j \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})$$

$$+ \quad \int_{2\pi} pdf(\omega_2)\frac{1}{pdf(\omega_2)}d\omega_2 \left[\Gamma_j(\boldsymbol{r_{Q_2}},\boldsymbol{u_2}) \int_{\sigma_2(\boldsymbol{r_{Q_2}},\boldsymbol{u_2})} pdf(\sigma_2)\frac{1}{pdf(\sigma_2)}d\sigma_2\right.$$

$$\times \quad \left.\left.\left.\kappa_i \tau_{ij} \kappa_j \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) + ... \quad \right]\right]\right] \tag{19}$$

which may be presented as [1]:

$$\varphi_{(V_i,V_j)} = \sum_{n=1}^{N_b} p_n \Delta\eta_n \int_0^1 pdf(g)dg \int_{V_i} pdf(V_i)dV_i$$

$$\int_{4\pi} pdf(\omega_0)d\omega_0 \int pdf(\sigma_0)d\sigma_0 \dots \int_{4\pi} pdf(\omega_m)d\omega_m \int pdf(\sigma_m)d\sigma_m \dots$$

$$\left[\quad W_{0,VV} + W_{1,VV} + W_{2,VV} + \dots + W_{m,VV} + \dots \quad\right] \tag{20}$$

where W_m represents the Monte Carlo weight corresponding to optical paths with m reflexions:

$$W_{0,VV} = \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \times \frac{\Gamma_j \kappa_i \tau_{ij} \kappa_j \Delta I_{ij}}{pdf(\sigma_0)} \tag{21}$$

$$W_{1,VV} = \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \times \frac{\Gamma_j \kappa_i \tau_{ij} \kappa_j \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})}{pdf(\sigma_1)pdf(\omega_1)} \tag{22}$$

$$\begin{aligned} W_{2,VV} &= \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \\ &\times \frac{\Gamma_j \kappa_i \tau_{ij} \kappa_j \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})}{pdf(\sigma_2)pdf(\omega_1)pdf(\omega_2)} \end{aligned} \tag{23}$$

---

[1] Noting that $\int_{4\pi} pdf(\omega_m)d\omega_m \equiv 1$ and $\int pdf(\sigma_m)d\sigma_m \equiv 1$, which allows to bring forward all integrals as an infinite product corresponding to all potential reflexion events.

$$W_{m,VV} = \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \tag{24}$$

$$\times \quad \frac{\Gamma_j\kappa_i\tau_{ij}\kappa_j\Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})\ldots\mathcal{F}_m(\boldsymbol{u_m}|\boldsymbol{u_{m-1}})}{pdf(\sigma_m)pdf(\omega_1)pdf(\omega_2)\ldots pdf(\omega_m)}$$

In the same way, we can write the expression of $\varphi_{(V_i,S_j)}$, $\varphi_{(V_i,S_j)}$, $\varphi_{(S_i,S_j)}$ (see appendix A).

Altogether, any optimized pdf set may be retained. The random sampling of each variable is performed according to this pdf set (narrow band, g, ω_0, σ_0, ω_1, σ_1, ω_2, σ_2, ...) leading to the definition of a multiple reflexion optical path and to the successive computations of all weights $W_0 \ldots W_m$ that are summed to obtain the total weight $W = W_1 + W_2 + \ldots W_m + \ldots$. This procedure is reproduced for a large number of sampled optical paths and the addressed radiative quantity is estimated as the average value of all W. Results presented in section 3 were obtained using the same optimized pdf set as in [15]. Statistical uncertainties are systematically associated to each result as for any Monte Carlo integration (see appendix C).

2.1.4 Truncation Error

The net radiative exchange expressions involve infinite sums over surface reflexions. Practically speaking a numerical truncation is therefore required. This truncation is performed here, as in standard MCM algorithms, on the basis of a user-defined required accuracy ϵ. The multiple reflexion optical path is constructed considering successive reflexions, until attenuation τ_{ij} is such that contributions of all further reflexions are lower than ϵ. The only specific point of MCM-NEF algorithms is that when following a ray, the successive values of ΔI_{ij} (the Planck function difference that appears in W_0, W_1, W_2,...) are only known when the second exchange point $\boldsymbol{r_{P_i}}$ is sampled. The truncation procedure therefore requires that an over-estimate of ΔI_{ij} is introduced, which may be taken as $I_\eta^0(T_{\max}) - I_\eta^0(T_{\min})$, where $T_{\max}$ and $T_{\min}$ are respectively the minimum and the maximum temperatures in the system.

2.2 Discrete Ordinates Method

Presently, the discrete ordinates method (DOM) [26, 27] is one of the most widely used radiation models. This is mainly due to the satisfactory accuracy of the model for most practical applications along with its flexibility, moderate computational requirements, and simplicity of coupling with CFD codes. The description of the method may be found in many publications, and therefore it is omitted here. However, the application of the DOM when the radiative properties of the medium are calculated using the CK method is much less common. This has been described in [3], and a short overview is given below for completeness.

In cylindrical coordinates (x, r, Ψ), the RTE for an emitting, absorbing and non-scattering medium may be written as:

$$\xi \frac{\partial I_\eta}{\partial x} + \frac{\beta}{r} \frac{\partial (r I_\eta)}{\partial r} - \frac{1}{r} \frac{\partial (\mu I_\eta)}{\partial \Psi} = -\kappa_\eta I_\eta + \kappa_\eta I_\eta^0 \tag{25}$$

where ξ, β and μ are the direction cosines.

The spatial and the angular discretizations of the RTE follow standard practices of the DOM. Hence, the spatial discretization was carried out using the finite volume method, yielding the following discretized equation for a direction m in the first quadrant (similar equations may be written for the other quadrants) and for a band of width $\Delta \eta_l$:

$$
\begin{aligned}
\xi^m & \left[I_{i+1/2,j,k,l}^m A_{i+1/2,j} - I_{i-1/2,j,k,l}^m A_{i-1/2,j} \right] \\
+ \quad & \beta^m \left[I_{i,j+1/2,k,l}^m A_{i,j+1/2} - I_{i,j-1/2,k,l}^m A_{i,j-1/2} \right] \\
- \quad & \left[A_{i,j+1/2} - A_{i,j-1/2} \right] \left[\frac{\alpha_{m+1/2} I_{i,j,k,l}^{m+1/2} - \alpha_{m-1/2} I_{i,j,k,l}^{m-1/2}}{w_m} \right] \\
= \quad & \left(-\kappa_{i,j,k,l} I_{i,j,k,l}^m + \kappa_{i,j,k,l} \overline{I}_{\eta,i,j,k,l}^0 \right) V
\end{aligned}
\tag{26}
$$

In this equation, the cell volume and the area of the cell faces were denoted by V and A, respectively, $\overline{I}_\eta^0$ is the mean blackbody radiation intensity over the band, w_m is the DOM quadrature weight and the coefficients α arise from the discretization of the third term on the left of Eq. 25, as explained, e.g., in [28]. The directions $m \pm 1/2$ define the edges of angle $\Delta \Psi$ associated with

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direction m. Subscripts i and j identify the control volume, and $i + 1/2$, $i - 1/2$, $j + 1/2$ and $j - 1/2$ refer to the east, west, north and south cell faces of that control volume, respectively. Subscripts k and l denote the CK quadrature point and the band under consideration. In this work, the cell face radiation intensities were calculated using the CLAM scheme [29], which is formally second-order accurate and bounded, and a level symmetric SN quadrature.

The total radiation intensity at a control volume is obtained by adding the contributions of all the CK quadrature points in a band, all directions and all bands:

$$I_{i,j} = \sum_{l=1}^{N_b} \Delta \eta_l \sum_{m=1}^{M} w_j \sum_{k=1}^{N_Q} \omega_k I_{i,j,k,l}^m \tag{27}$$

where ω_k is the CK quadrature weight, N_b the number of bands, M the number of directions, and N_Q the number of quadrature points in the CK method. There are 7 quadrature points per band and per participating gas; the weighting factors ω_k are taken from the database of Ecole Centrale de Paris [30, 25].

The heat flux incident on the west boundary is calculated as:

$$q_{w,j} = \sum_{l=1}^{N_b} \Delta \eta_l \left[\sum_{k=1}^{N_Q} \omega_k \left(\sum_{\substack{m=1 \\ (\xi^m < 0)}}^{M} w_m |\xi^m| I_{i,j,k,l}^m \right) \right] \tag{28}$$

Similar equations may be written for the other boundaries. The divergence of the radiative heat flux is obtained from:

$$\nabla . q_{i,j} = \sum_{l=1}^{N_b} \Delta \eta_l \left[\sum_{k=1}^{N_Q} \omega_k k_{i,j,k,l} \left(4\pi \overline{I}_{\eta,i,j,k,l}^0 - \sum_{m=1}^{M} w_m I_{i,j,k,l}^m \right) \right] \tag{29}$$

3   Results and Discussion

3.1   Description of the test cases

Three radiative heat transfer problems in two-dimensional axisymmetric enclosures involving gas-soot mixtures at atmospheric pressure and gray diffuse reflective walls have been solved. Specular walls have also been introduced and comparisons with diffuse walls calculations have been

performed. In the first test case, a homogeneous, isothermal configuration is considered and both the radiative heat source along the centerline of the cylinder and the incident radiative heat flux on the lateral wall are computed. Test case 2 presents a homogeneous, non-isothermal configuration. The radiative heat source is calculated along the centerline of the cylinder and the net wall heat flux is calculated on the lateral wall. Spectral data for soot and gas are the same as those used in [3]. As soot particle diameters are nanometric, the scattering effect is neglected. A third test case is considered to highlight the effect of considering either diffuse or specular walls.

The level of accuracy of the CK compared to the SNB model has been discussed in detail in our previous publication [3]; errors due to the use of these two different models do not exceed 1%. Comparisons between the SNB-CK and the SNB models have already been studied in previous papers [31, 32, 33, 34]. In the configurations that we present here (test case 1 and 2), the SNB-CK and the SNB give solutions in very close agreement; the error between computed solutions by both methods (DOM and MCM) is mainly due to the angular and spatial discretizations of the RTE case is therefore ideally suited to highlight the difference between specular and diffuse walls.

3.1.1 Test case 1

The enclosure is a cylinder of length $L = 3m$ and radius $R = 0.5m$. The walls are gray with an emissivity $\varepsilon_w = 0.5$, and a temperature of 800K. The temperature of the gas is 1800K. The composition of the medium is 20 % H_2O, 10 % CO_2 and 70 % N_2 at atmospheric pressure and the soot volumetric fraction is $f_v = 10^{-7}$. The radiative heat source along the centerline of the cylinder is shown in Figures 3(a) and 3(b). The incident heat flux on the circular wall is shown in Figures 3(c) and 3(d). Both the MCM and the DOM results are shown. The DOM calculations have been carried out for two different grids, respectively with 41×30 (coarse grid) and 81×60 control volumes. Two different quadratures were also employed: $S8$ and $S16$.

The radiative heat source presents a strong variation for calculation points that are very close to the cold side walls ($x = 0$ and $x = L$). Between $x \gtrsim 0.5m$ and $x \lesssim 2.5m$, the profile is roughly flat with a heat source of $\approx 500kW.m^{-3}$ at the center of the cylinder. The incident heat flux presents also a quite flat profile far from the side walls and decreases with the distance to these

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walls. Figure 3 shows a minor influence of the spatial discretization on the DOM results in this test case because the medium is homogeneous and isothermal. The accuracy of the DOM is improved with the refinement of the angular discretization. As an indication, the difference between DOM and MCM results was compared at the center point ($x = 1.5m$) of the cylinder. For DOM-S16 calculations, the heat source was found to be 1.38% less than the MCM results for the coarse grid (2.40% for DOM-S8), and 1.34% less for the finer one (2.19% for DOM-S8). Considering the wall heat flux, a difference of 0.32% was found for DOM-S16 results with the coarse grid (1.31% for DOM-S8) compared to MCM ones, and a difference of 0.31% for the finer one (1.31% for DOM-S8).

The numerical results associated to the MCM calculations can be found in appendix B (Table 1) in order to enable readers to use them for benchmark purposes. Results of DOM calculations for specular walls are not available on this configuration, but a comparison between specular and diffuse walls is presented, using MCM calculations on figure 4.

3.1.2   Test case 2

An axisymmetric enclosure is considered again where the length of the cylinder is $L = 1.2m$ and the radius is $R = 0.3m$. The medium is homogeneous, composed of 15% water vapor and 85% nitrogen. The soot volumetric fraction is $f_v = 10^{-6}$. The lateral wall is diffuse with an emissivity of $\varepsilon_w = 0.8$; the side walls ($x = 0m$ and $x = L = 1.2m$) are black. The following two-dimensional gas temperature profile is considered:

$$T(x,r) = 800 + 1200\left(1 - \frac{r}{R}\right)\left(\frac{x}{L}\right) \tag{30}$$

This means that all the walls have a temperature of $800K$ except the right wall ($x = L$), which is maintained at $300K$.

The divergence of the radiative heat flux at the centerline of the cylinder is continuously increasing with x, because of the temperature profile of the medium and the surrounding walls, which are maintained at the same temperature. Both the source term and the gas temperature reach the maximum at the immediate vicinity of the cold wall ($x = L$) (Figures 5(a) and 5(b)). Figures

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5(c) and 5(d) show that the net heat flux is also increasing continuously from $x = 0m$ to $x \approx 1m$, and then decreases because of the presence of the cold wall.

As for the test case 1 the calculations are carried out using the DOM for two grids and two angular discretizations and are compared to the MCM results (Figure 5). The DOM results are in excellent agreement with the MCM results for the calculation of the radiative heat source on the centerline of the cylinder; the refinement of the grid seems to have no influence on the results, and the angular discretization has a minor influence. In Figure 5(c), a discrepancy is observed between the DOM-S8 and the MCM results for the calculation of the net heat flux on the lateral wall. The discrepancy is not observed any more with a finer angular discretization, which suggests that the difference was only due to the ray effect. The MCM numerical results for test case 2 can be found in appendix B (Table 2).The DOM (coarse grid) numerical results are available too in the appendix B (Table 3) at some calculation points. The radiative heat source term and the net heat fluxes are computed by the MCM for diffuse and specular walls and are compared in figure 6; the difference between both solution remains marginal.

3.1.3   Test case 3

The geometry (see figure 7) is axisymmetric with a length $L = 0.24m$ and a radius $0.03m$. The whole configuration is at a temperature of $T = 1500K$, except the top of the cylinder at $x = 0.24m$, which is at $400K$. The medium is a mixture of 20% H_2O, 10% of CO_2 and a volumetric fraction of soot $f_v = 10^{-7}$. The walls are black, except the bottom third of the cylindrical wall (from $x = 0m$ to $x = 0.08m$), which is taken either as black, diffuse or specular ($\varepsilon = 0.5$). The average value of the heat source is computed within a cylinder of same height (between $x = 0m$ and $x = 0.08m$) and variable radius (between $r = 0m$ and $r = 0.03m$) to show the influence of the specular reflexion (Figure 8). For black and diffuse walls, the average radiative source term is quasi independent of the radius (within the error bars).When the radius is equal to 0.03 (radius of the cylindrical enclosure), the difference between the diffuse and specular reflexion is about 5%, an order of magnitude that is comparable to the differences observed in the preceeding test cases. When the radius decreases, this difference increases and reaches 20%. This is due to the fact that

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the reflective surfaces are restricted to the bottom third of the cylindrical surface : with specular properties, no optical path can be found to reach the top cold surface with at least one reflexion. The average heat source is therefore due to the direct optical paths (without reflexion) only and indeed for small radius, this average radiative source term tends to the value observed for black surfaces. This test case is therefore ideally suited to highlight the difference between specular and diffuse walls.

4   Conclusion

A Monte Carlo Net Exchange Formulation including multiple reflexions has been detailed and used to solve two radiative heat transfer problems. For all test cases, diffuse and specular reflexions were considered with a medium that is a mixture of H_2O, CO_2, N_2 and soot, at atmospheric pressure. A Discrete Ordinates Method has also been used on the two first test cases in order to compare with the Monte Carlo results in the case of diffuse reflexion. Altogether :

- The MCM-NEF has been extended to multidimensional enclosures with reflective, diffuse or specular boundaries.

- Highly accurate solutions are available for the three considered configurations that can be used for benchmarking.

- The DOM-CK method was shown to perform satisfactorily using the S8 quadrature (maximum error of 2.4% for radiative source term and 1.31% for radiative wall heat flux) and to yield very good predictions using S16 quadrature (maximum error of 1.38% for radiative source term and 0.32% for radiative wall heat flux).

A Weights for volume-surface, surface-volume, and surface-surface exchanges

$$
\begin{aligned}
\varphi_{(V_i,S_j)} \quad = \quad & \sum_{n=1}^{N_b} p_n \Delta_{\nu n} \int_0^1 pdf(g)dg \int_{V_i} pdf(V_i)dV_i \\
& \int_{4\pi} pdf(\omega_0)d\omega_0 \ldots \int_{4\pi} pdf(\omega_m)d\omega_m \ldots \\
& \left[W_{0,VS} + W_{1,VS} + W_{2,VS} + \ldots + W_{m,VS} + \ldots \right]
\end{aligned}
\tag{31}
$$

$$
W_{0,VS} = \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \times \frac{\Gamma_j \kappa_i \tau_{ij} \varepsilon_1 \Delta I_{ij}}{1}
\tag{32}
$$

$$
W_{1,VS} = \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \times \frac{\Gamma_j \kappa_i \tau_{ij} \varepsilon_2 \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})}{pdf(\omega_1)}
\tag{33}
$$

$$
\begin{aligned}
W_{2,VS} \quad = \quad & \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \\
\times \quad & \frac{\Gamma_j \kappa_i \tau_{ij} \varepsilon_3 \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})}{pdf(\omega_1)pdf(\omega_2)}
\end{aligned}
\tag{34}
$$

$$
\begin{aligned}
W_{m,VS} \quad = \quad & \frac{1}{p_n \quad pdf(g)pdf(V_i)pdf(\omega_0)} \\
\times \quad & \frac{(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})\ldots(1-\varepsilon_m)\mathcal{F}_m(\boldsymbol{u_m}|\boldsymbol{u_{m-1}})}{pdf(\omega_1)pdf(\omega_2)\ldots pdf(\omega_m)} \\
\times \quad & \Gamma_j \kappa_i \tau_{ij} \varepsilon_{m+1} \Delta I_{ij}
\end{aligned}
\tag{35}
$$

$$\varphi_{(S_i, V_j)} = \sum_{n=1}^{N_b} p_n \Delta_{\nu n} \int_0^1 pdf(g) d\kappa \int_{S_i} pdf(S_i) dS_i$$

$$\int_{4\pi} pdf(\omega_0) d\omega_0 \int pdf(\sigma_0) d\sigma_0 \ldots$$

$$\int_{4\pi} pdf(\omega_m) d\omega_m \int pdf(\sigma_m) d\sigma_m \ldots$$

$$\left[W_{0,SV} + W_{1,SV} + W_{2,SV} + \ldots + W_{m,SV} + \ldots \right] \tag{36}$$

$$W_{0,SV} = \frac{1}{p_n \quad pdf(g) pdf(S_i) pdf(\omega_0)} \times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij}}{pdf(\sigma_0)} \tag{37}$$

$$W_{1,SV} = \frac{1}{p_n \quad pdf(g) pdf(S_i) pdf(\omega_0)} \times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})}{pdf(\sigma_1) pdf(\omega_1)} \tag{38}$$

$$
\begin{aligned}
W_{2,SV} &= \frac{1}{p_n \quad pdf(g) pdf(S_i) pdf(\omega_0)} \\
&\times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij} (1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})}{pdf(\sigma_2) pdf(\omega_1) pdf(\omega_2)}
\end{aligned}
\tag{39}
$$

$$
\begin{aligned}
W_{m,SV} &= \frac{1}{p_n \quad pdf(g) pdf(S_i) pdf(\omega_0)} \\
&\times \frac{(1 - \varepsilon_1) \mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1 - \varepsilon_2) \mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1}) \ldots (1 - \varepsilon_m) \mathcal{F}_m(\boldsymbol{u_m}|\boldsymbol{u_{m-1}})}{pdf(\sigma_m) pdf(\omega_1) pdf(\omega_2) \ldots pdf(\omega_m)} \\
&\times \Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \kappa_j \Delta I_{ij}
\end{aligned}
\tag{40}
$$

$$\varphi_{(S_i,S_j)} = \sum_{n=1}^{N_b} p_n \Delta_{\nu n} \int_0^1 pdf(g) d\kappa \int_{S_i} pdf(S_i) dS_i$$

$$\int_{4\pi} pdf(\omega_0) d\omega_0 \ldots \int_{4\pi} pdf(\omega_m) d\omega_m \ldots$$

$$\left[W_{0,SS} + W_{1,SS} + W_{2,SS} + \ldots + W_{m,SS} + \ldots \right] \tag{41}$$

$$W_{0,SS} = \frac{1}{p_n \quad pdf(g)pdf(S_i)pdf(\omega_0)} \times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_1 \Delta I_{ij}}{1} \tag{42}$$

$$W_{1,SS} = \frac{1}{p_n \quad pdf(g)pdf(S_i)pdf(\omega_0)} \times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_2 \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})}{pdf(\omega_1)} \tag{43}$$

$$
\begin{aligned}
W_{2,SS} &= \frac{1}{p_n \quad pdf(g)pdf(S_i)pdf(\omega_0)} \\
&\times \frac{\Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_3 \Delta I_{ij}(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})}{pdf(\omega_1)pdf(\omega_2)}
\end{aligned} \tag{44}
$$

$$
\begin{aligned}
W_{m,SS} &= \frac{f(\kappa)}{p_n \quad pdf(\kappa)pdf(S_i)pdf(\omega_0)} \\
&\times \frac{(1-\varepsilon_1)\mathcal{F}_1(\boldsymbol{u_1}|\boldsymbol{u_0})(1-\varepsilon_2)\mathcal{F}_2(\boldsymbol{u_2}|\boldsymbol{u_1})\ldots(1-\varepsilon_m)\mathcal{F}_m(\boldsymbol{u_m}|\boldsymbol{u_{m-1}})}{pdf(\omega_1)pdf(\omega_2)\ldots pdf(\omega_m)} \\
&\times \Gamma_j \varepsilon_i \cos\theta_i \tau_{ij} \varepsilon_{m+1} \Delta I_{ij}
\end{aligned} \tag{45}
$$

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B   Numerical results for test case 1 and test case 2

Test case 1					
x (m)	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$	x (m)	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$
0.0205	770.83±1.58	232.07±0.39	0.6505	513.84±1.54	300.06±0.48
0.0405	697.80±1.47	242.02±0.41	0.7005	511.13±1.53	300.84±0.51
0.0605	657.24±1.45	249.33±0.41	0.7505	509.64±1.55	302.39±0.52
0.0805	632.61±1.45	254.26±0.40	0.8005	509.40±1.51	302.18±0.51
0.1005	614.27±1.41	259.22±0.42	0.8505	506.26±1.48	303.88±0.51
0.1205	601.34±1.38	262.94±0.37	0.9005	505.15±1.45	304.63±0.53
0.1405	589.66±1.52	266.72±0.44	0.9505	503.50±1.45	305.38±0.49
0.1605	582.51±1.53	269.55±0.47	1.0005	502.54±1.49	305.67±0.52
0.1805	575.40±1.59	272.38±0.46	1.0505	501.63±1.49	306.21±0.55
0.2005	568.26±1.53	274.10±0.48	1.1005	501.46±1.48	306.69±0.53
0.2505	555.22±1.47	279.82±0.45	1.1505	501.01±1.56	307.34±0.52
0.3005	547.66±1.47	283.67±0.42	1.2005	499.73±1.45	307.88±0.51
0.3505	537.50±1.57	287.61±0.47	1.2505	499.61±1.64	307.85±0.49
0.4005	531.34±1.55	290.29±0.49	1.3005	499.44±1.64	307.10±0.51
0.4505	528.04±1.57	292.47±0.51	1.3505	499.51±1.63	308.10±0.54
0.5005	523.93±1.56	295.15±0.51	1.4005	499.56±1.77	308.39±0.53
0.5505	519.16±1.51	296.60±0.51	1.4505	496.62±1.64	308.12±0.52
0.6005	516.65±1.56	298.08±0.47	1.4905	498.89±1.59	307.91±0.49

Table 1: Numerical results for the first test case; q_w is the flux on the lateral wall of the cylinder, and $\nabla.\mathbf{q}$ is the radiative source term on the axis of the cylinder. Data from the MCM and associated statistical errors.

Test case 2					
x (m)	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$	x (m)	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$
0.025	-5.25±1.05	2.20±0.03	0.625	1281.18±11.63	20.15±0.15
0.075	29.63±1.34	2.98±0.03	0.675	1539.15±13.54	22.36±0.17
0.125	69.98±1.92	3.92±0.04	0.725	1827.86±14.24	24.54±0.18
0.175	119.22±2.45	4.97±0.05	0.775	2162.77±15.56	26.81±0.20
0.225	177.76±3.40	6.17±0.06	0.825	2554.37±17.50	28.94±0.22
0.275	250.26±3.75	7.50±0.07	0.875	2990.06±18.55	30.85±0.23
0.325	335.50±4.64	8.98±0.08	0.925	3489.83±20.37	32.47±0.25
0.375	435.91±5.95	10.54±0.09	0.975	4065.79±22.18	33.20±0.26
0.425	558.42±6.72	12.24±0.10	1.025	4741.95±24.85	33.07±0.26
0.475	699.10±7.44	14.06±0.11	1.075	5549.90±28.30	31.56±0.27
0.525	864.32±8.50	16.05±0.12	1.125	6530.07±31.34	27.80±0.95
0.575	1062.02±10.19	18.02±0.14	1.175	7992.95±3.18	20.70±1.16

Table 2: Numerical results for the first test case; q_w is the net flux on the lateral wall of the cylinder, and $\nabla.\mathbf{q}$ is the radiative source term on the axis of the cylinder. Data from the MCM and associated statistical errors.

x (m)	Test case 2: DOM-$S16$		Test case 2: DOM-S8	
	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$	$\nabla.\mathbf{q}(kW/m^3)$	$q_w(kW/m^2)$
0.075	28.28	3.03	19.19	2.67
0.225	185.67	6.20	171.26	5.73
0.375	438.25	10.48	435.96	9.9
0.525	901.89	15.83	868.35	15.24
0.675	1587.00	22.03	1540.04	21.61
0.825	2604.75	28.46	2544.99	28.02
0.975	4090.58	32.62	4017.71	30.93
1.125	6404.20	27.88	6328.39	28.17

Table 3: Numerical results for the second test case; q_w is the net flux on the lateral wall of the cylinder, and $\nabla.\mathbf{q}$ is the radiative source term on the axis of the cylinder. Results from the DOM calculations.

C Monte Carlo integration principle and statistical errors

The statistical errors appearing in tables B.1 and B.2 are computed together with the solution according to the standard Monte Carlo integration method which principle is reminded hereafter.

Let us consider the integral $I = \int_D f(x)dx$. One can always introduce an arbitrary probability density function p defined and strictly positive on the integration domain D and define $g(x) = \frac{f(x)}{p(x)}$ so that:

$$I = \int_D \frac{f(x)}{p(x)}p(x)dx = \int_D g(x)p(x)dx$$

Let us now define a random variable X distributed according to p, then $g(X)$ is also a random variable and I is the expectation of $g(X)$, and I will be estimated with N samples of $g(X)$:

$$I = E[g(X)] \approx \frac{1}{N}\sum_{i=1}^{N} g(x_i) = <g(X)>_N \quad \left(I = \lim_{N\to\infty} <g(X)>_N \right)$$

The standard deviation of the estimate is $\sigma(<g(X)>_N) = \frac{1}{\sqrt{N}}\sigma(g(X))$, where $\sigma(g(X))$ is the standard deviation of $g(X)$, and will be approximated as:

$$\sigma(<g(X)>_N) \approx \frac{1}{\sqrt{N}}\sqrt{[<g(X)^2>_N - <g(X)>_N^2]}$$

The relative errors reported in tables B.1 and B.2 correspond to this standard deviation estimate divided by the integral estimate.

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Figure captions

Figure 1:

Exchange between two volumes in case of black enclosures.

Figure 2:

Exchange between two volumes in case of reflective walls.

Figure 3:

Heat source on the axis of a cylinder and radiative flux on the lateral side for test case 1.

Figure 4:

Comparison between a specular and a diffuse wall for test case 1.

Figure 5:

Heat source on the axis of a cylinder and radiative net flux on the lateral side for test case 2.

Figure 6:

Comparison between a specular and a diffuse wall for test case 2.

Figure 7:

Average heat source term $\frac{\int_C -\boldsymbol{\nabla}.\boldsymbol{q}_r}{\pi r^2 \frac{L}{3}}$ computed for a cylinder with variable radius within a cylindrical enclosure

Figure 8:

Schematic of test case 3.
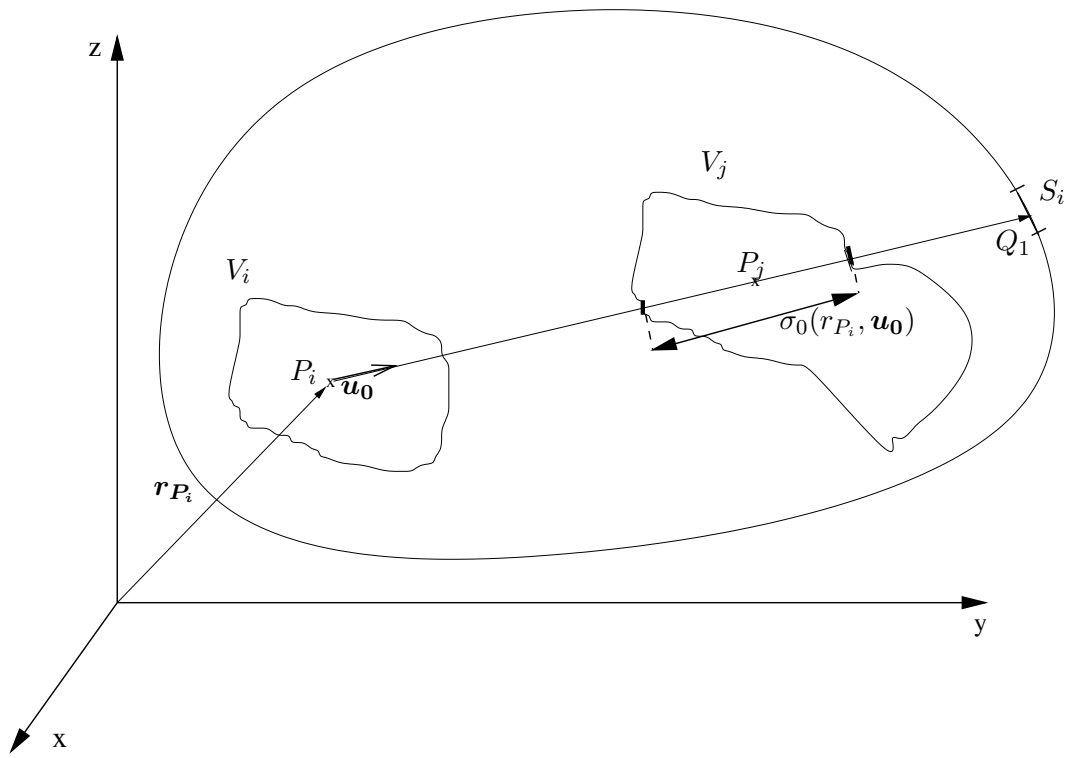
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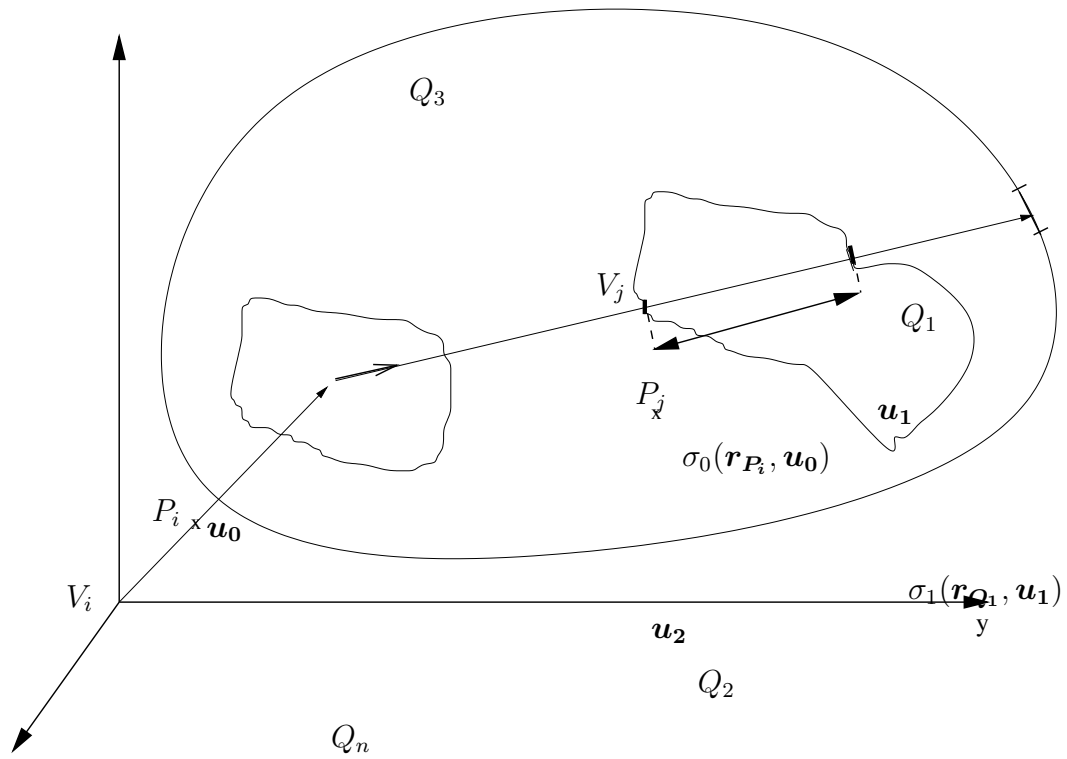
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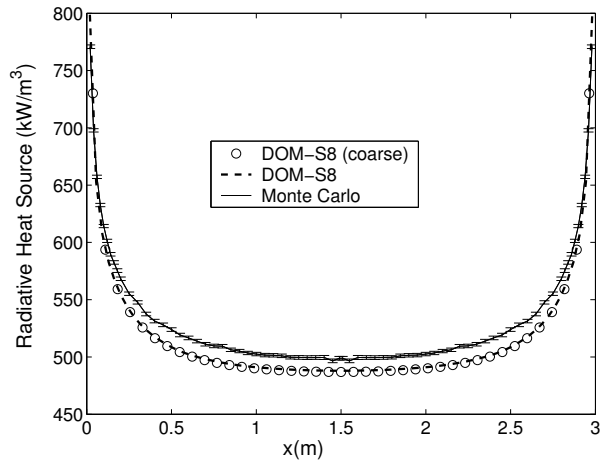
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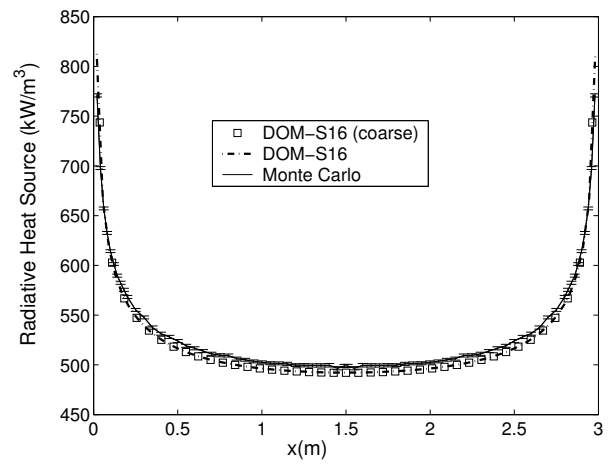
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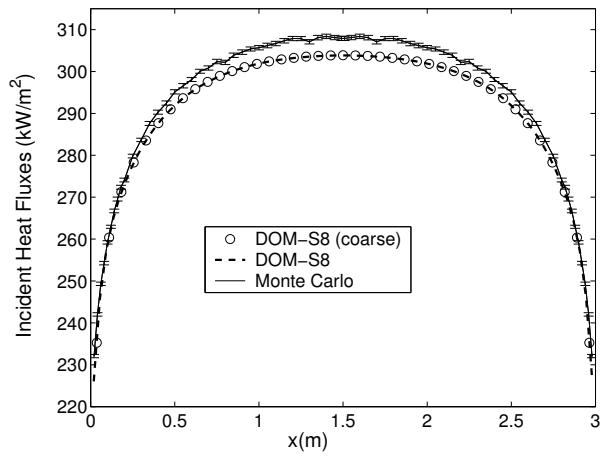
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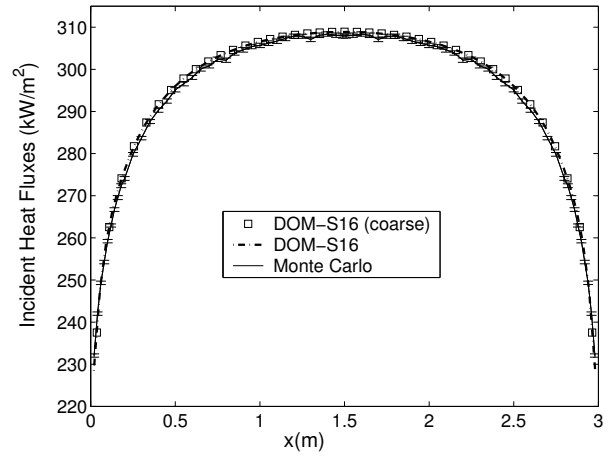
(d)

Figure 3:

(a)

(b)

Figure 4:

36
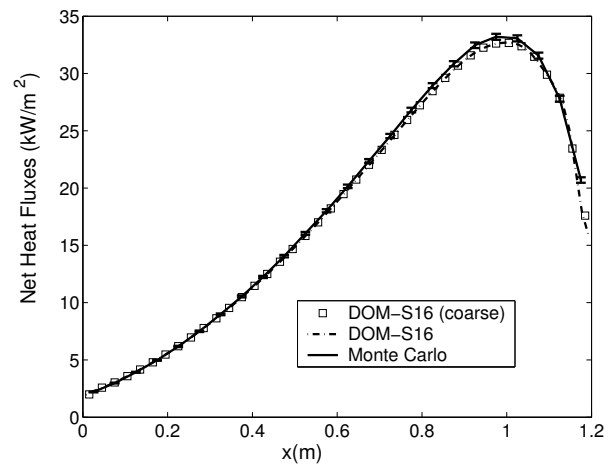
(a)   (b)

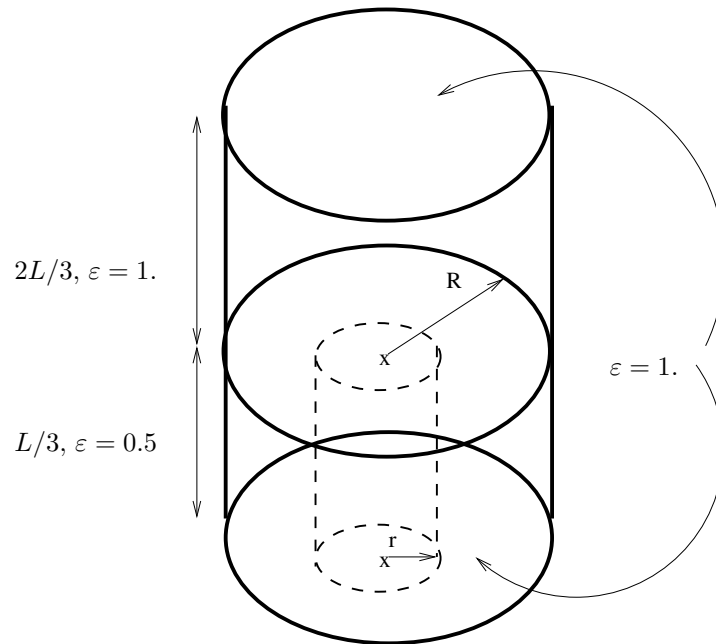(c)   (d)
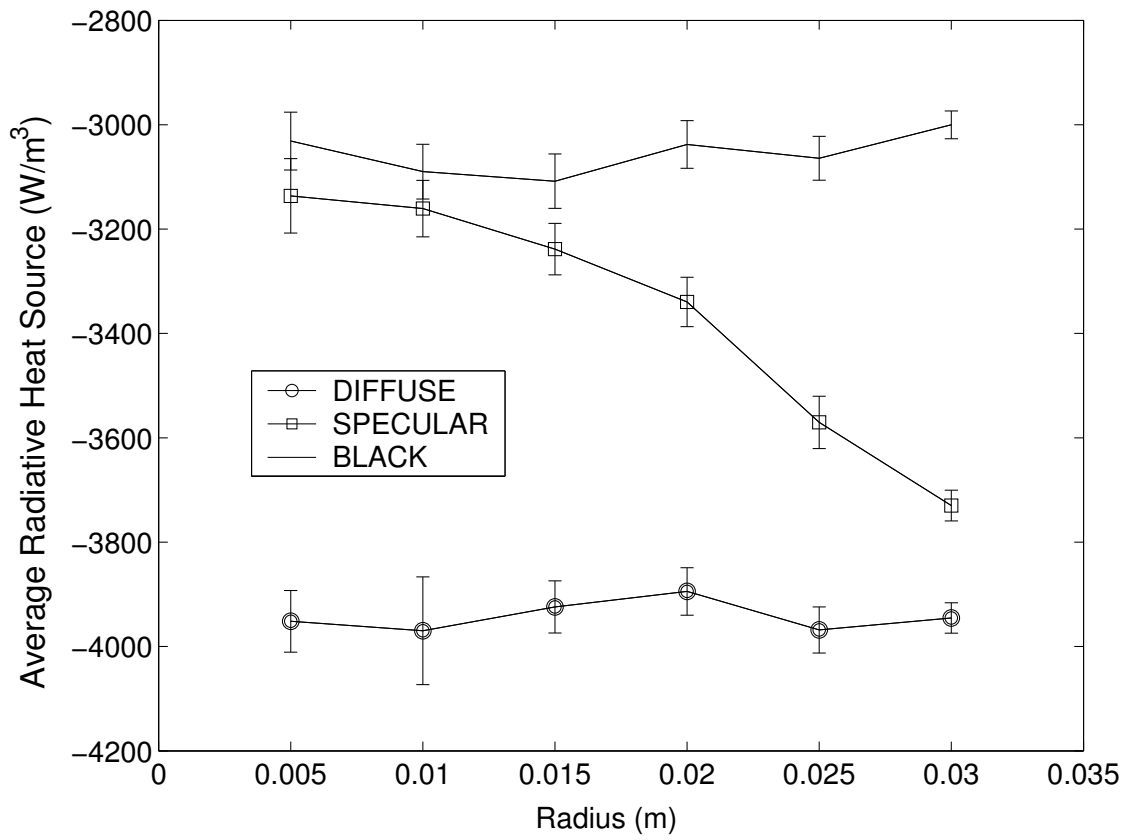
Figure 5:

(a)

(b)

Figure 6:

$2L/3,\ \varepsilon = 1.$

R

$\varepsilon = 1.$

$L/3,\ \varepsilon = 0.5$

r

x

x

Figure 7:

Figure 8: